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Daily Rainfall Estimate by Emissivity Temporal Variation from 10 Satellites --Manuscript Draft--

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Abstract:	Rainfall retrieval algorithms for passive microwave radiometers often exploits the brightness temperature depression due to ice scattering at high frequency channels (\geq 85 GHz) over land. This study presents an alternate method to estimate the daily rainfall amount using the emissivity temporal variation (i.e., Δe) under rain-free conditions at low frequency channels (19, 24 and 37 GHz). Emissivity is derived from 10 passive microwave radiometers, including the Global Precipitation Measurement (GPM) Microwave Imager (GMI), the Advanced Microwave Scanning Radiometer 2 (AMSR2), three Special Sensor Microwave Imager/Sounder (SSMIS), the Advanced Technology Microwave Sounder (ATMS), and four Advanced Microwave Sounding Unit-A (AMSU-A). Four different satellite combination schemes are used to derive the Δe for daily rainfall estimates. They are all-10-satellites, 5-imagers, 6-satellites with very different equator crossing times, and GMI-only. Results show that Δe from all-10-satellites has the best performance with a correlation of 0.60 and RMSE of 6.52 mm, comparing with the integrated multi-satellite retrievals (IMERG) final run product. The 6-satellites scheme has comparable performance with all-10-satellites scheme. The 5-imagers scheme performs noticeably worse with a correlation of 0.49 and RMSE of 7.28 mm, while the GMI-only scheme performance from the 5-imagers and GMI-only schemes can be explained by the much longer revisit time, which cannot accurately capture the emissivity temporal variation.			
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Dear Editor,

We would like to submit our manuscript entitled "Daily Rainfall Estimate by Emissivity Temporal Variation from 10 Satellites" for consideration of publication in the Journal of Hydrometeorology. This study presents an alternate method to estimate the daily rainfall amount using the emissivity temporal variation (i.e., Δe) under rain-free conditions at low frequency channels (19, 24 and 37 GHz) from 10 satellites. The key innovation is that we apply the "soil-moisture-based" retrieval concept to the low frequency channels from 10 satellites. It is of relevance to the broader field of remote sensing of precipitation and hydrometeorology. Therefore, we believe that our article falls well within the scope of the Journal of Hydrometeorology, and should be of interest to the journal's core readership.

We confirm that this manuscript has not been published elsewhere and is not under consideration by another journal. All authors have approved the manuscript and agree with its submission. Thank you for your assistance in handling the review of this manuscript.

Sincerely,

Yalei You

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ABSTRACT

Rainfall retrieval algorithms for passive microwave radiometers often ex-14 ploits the brightness temperature depression due to ice scattering at high fre-15 quency channels (\geq 85 GHz) over land. This study presents an alternate 16 method to estimate the daily rainfall amount using the emissivity temporal 17 variation (i.e., Δe) under rain-free conditions at low frequency channels (19, 18 24 and 37 GHz). Emissivity is derived from 10 passive microwave radiome-19 ters, including the Global Precipitation Measurement (GPM) Microwave Im-20 ager (GMI), the Advanced Microwave Scanning Radiometer 2 (AMSR2), 21 three Special Sensor Microwave Imager/Sounder (SSMIS), the Advanced 22 Technology Microwave Sounder (ATMS), and four Advanced Microwave 23 Sounding Unit-A (AMSU-A). Four different satellite combination schemes 24 are used to derive the Δe for daily rainfall estimates. They are all-10-satellites, 25 5-imagers, 6-satellites with very different equator crossing times, and GMI-26 only. Results show that Δe from all-10-satellites has the best performance with 27 a correlation of 0.60 and RMSE of 6.52 mm, comparing with the integrated 28 multi-satellite retrievals (IMERG) final run product. The 6-satellites scheme 29 has comparable performance with all-10-satellites scheme. The 5-imagers 30 scheme performs noticeably worse with a correlation of 0.49 and RMSE of 31 7.28 mm, while the GMI-only scheme performs the worst with a correlation of 32 0.25 and RMSE of 11.36 mm. The inferior performance from the 5-imagers 33 and GMI-only schemes can be explained by the much longer revisit time, 34 which cannot accurately capture the emissivity temporal variation. 35

36 1. Introduction

Spaceborne passive microwave radiometers have long been recognized as key instruments for 37 global rainfall estimates over land. Early studies in 1980s from the Scanning Multichannel Mi-38 crowave Radiometer (SSMR) onboard Nimbus-7 satellite showed that brightness temperature 39 (TB) has a clear depression signature under thunderstorms due to the ice particles' scattering 40 effect in the atmosphere (Spencer et al. 1983a,b). Many satellites after Nimbus-7 carried the ra-41 diometers capable of estimating the surface rainfall rate over land via this ice scattering concept. 42 These radiometers include Special Sensor Microwave Imager (SSMI), Special Sensor Microwave 43 Imager/Sounder (SSMIS), Tropical Rainfall Measuring Mission Microwave Imager (TMI), Ad-44 vanced Microwave Scanning Radiometer—Earth Observing System (AMSR-E) and its successor, 45 AMSR2, Global Precipitation Measurement (GPM) Microwave Imager (GMI), MicroWave Ra-46 diation Imager (MWRI), Advanced Microwave Sounding Unit-A and B (AMSU-A/B), and Ad-47 vanced Technology Microwave Sounder (ATMS). Rainfall estimates from these radiometers serve 48 as the backbone for generating the widely-used global precipitation datasets, including NASA's 49 integrated multi-satellite retrievals (IMERG) (Huffman et al. 2015), Climate Prediction Center's 50 morphing technique (CMORPH) (Xie et al. 2017), and JAXA's Global Satellite Mapping of Pre-51 cipitation (GSMaP) dataset (Kubota et al. 2007). 52

Rainfall retrieval algorithm development has been extensively researched for these passive microwave radiometers. For example, Spencer et al. (1989) proposed the "Polarization Corrected Temperature" (PCT) to detect and retrieve rainfall over land from SSMI. Grody (1991) developed the "scattering index" (SI) technique to estimate the rainfall over land for SSMI, which is later improved by Ferraro et al. (1994); Ferraro and Marks (1995). Both PCT and SI methods primarily rely on the TB observations at 85 GHz over land, which is the highest available frequency

on SSMI. With the successful launch of the TRMM satellite in 1997, many retrieval algorithms for precipitation over land have also been developed for TMI (Wang et al. 2009; Gopalan et al. 60 2010; Petty and Li 2013b; Islam et al. 2014). In addition, retrieval algorithms have also been 61 developed for sensors with even higher frequencies (e.g., 150 and 183 GHz), including SSMIS 62 (You et al. 2015) and ATMS (Surussavadee and Staelin 2010; You et al. 2015). Different from 63 these sensor-specific algorithms, generic retrieval algorithms with the capability of adapting to 64 all these radiometers were also developed, including the Goddard profiling algorithm (GPROF) 65 (Kummerow et al. 2015), the Microwave Integrated Retrieval System (MiRS) (Boukabara et al. 66 2011), GSMaP level-2 precipitation retrieval algorithm (Aonashi et al. 2009; Shige et al. 2009), 67 and the one-dimensional variational (1DVAR) retrieval model (Meng et al. 2017). 68

To more accurately estimate the surface rainfall rate, these "hydrometeor-based" algorithms 69 have to mitigate the influence from the land surface, since TB reflects the integrated effect from 70 the hydrometeors in the air and the surface emission. To this end, the entire globe is divided into 5 71 by 5 degree grid boxes in each season in the GSMaP retrieval process (Aonashi et al. 2009). Simi-72 larly, ancillary land surface type information (Aires et al. 2011) has been used by several retrieval 73 algorithms (You et al. 2015; Kummerow et al. 2015). To largely avoid the possible surface con-74 tamination, instead of using the signatures from window channels (e.g., 85GHz), Staelin and Chen 75 (2000) developed a rainfall retrieval algorithm solely dependent on the microwave observations 76 near opaque water vapor and oxygen absorption channels (183 GHz and 52 GHz). 77

⁷⁸ Brocca et al. (2014) proposed a conceptually different rainfall retrieval algorithm by using the ⁷⁹ soil moisture datasets derived from spaceborne microwave sensors. They concluded that the re-⁸⁰ trieved 5-day rainfall accumulation from the soil moisture datasets agree reasonably well with a ⁸¹ ground gauge analysis dataset, indicated by the correlation being as large as 0.54. The ability ⁸² to retrieve rainfall from the soil moisture is further demonstrated by Koster et al. (2016), which showed that satellite missions designed for soil moisture observations indeed contain valuable
rainfall information. In fact, soil moisture information has also been exploited to improve the
hydrometeor-based rainfall retrieval results (Crow et al. 2009; Pellarin et al. 2013).

There are key differences between these two types of rainfall algorithms, which are referred to 86 as "hydrometeor-based" and "soil-moisture-based" retrieval algorithms for convenience. First, the 87 microwave sensors designed for soil moisture measurement utilize lower frequencies than those 88 suitable for hydrometeor measurement. For example, the radiometers onboard the Soil Moisture 89 Active-Passive (SMAP) satellite and the Soil Moisture and Ocean Salinity (SMOS) satellite have a 90 frequency of 1.4 GHz. The Advanced Scatterometer (ASCAT) onboard the MetOp satellites oper-91 ates at \sim 5.2 GHz. In contrast, the primary frequencies to measure the ice scattering over land from 92 passive microwave radiometers are around 85 GHz and higher (e.g., 150 and 183 GHz). The lower 93 frequencies used for soil moisture measurement can penetrate a thicker layer of soil and thereby provide more information about the rainfall impact on the soil, while the higher frequencies are 95 more sensitive to the hydrometeors in the atmosphere. Second, the hydrometeor-based algorithm 96 attempts to minimize the possible surface contamination (e.g., soil moisture, surface temperature, 97 and vegetation). On the contrary, the soil-moisture-based algorithm attempts to limit the impact 98 from the hydrometeors. Third, the hydrometeor-based algorithm uses the instantaneous observations at the time of the overpass, providing a snapshot of the rainfall rate at that time. In contrast, 100 the soil-moisture-based algorithms use observations that are not contaminated by hydrometeors 101 in the atmosphere, and therefore are more representative of accumulated rainfall over some time 102 prior to the observation. 103

The objective of this study is to estimate the daily rainfall accumulation from the land surface emissivity variation due to the rainfall impact. Previous work showed that the land surface microwave emissivity tends to decrease after rainfall due to the increase of soil moisture (Jackson

1993; Ferraro et al. 2013; You et al. 2014; Yin et al. 2019). In other words, the land surface mi-107 crowave emissivity variation is directly related to soil moisture change. Therefore, in essence, this 108 work attempts to relate the soil moisture variation to the rainfall accumulation, similar to Brocca 109 et al. (2014) and Koster et al. (2016). The key innovation is that we apply the soil-moisture-based 110 retrieval concept to the low frequency channels (19, 24 and 37 GHz) from 10 satellites (Table 1), 111 instead of soil moisture-specific channels (1.4 GHz) that are only available on one or two satellites. 112 Previous rainfall retrieval algorithms for these 10 sensors estimated the instantaneous rainfall 113 rate based on the ice scattering signal primarily from the high frequencies (>85 GHz) (e.g., Ferraro 114 et al. 2000; Ebtehaj et al. 2015; You et al. 2015; Kummerow et al. 2015; You et al. 2016). For the 115 lower frequency channels, the ice scattering signal is less pronounced over land due to the longer 116 wavelength (Spencer et al. 1983a). In addition, the high and highly variable land surface emissivity 117 often masks out the liquid raindrop emission signal at the low frequency channels (Prigent et al. 118 2006; Munchak et al. 2020). For these reasons, these low frequency channels are either not used or 119 play a secondary role in the instantaneous rainfall retrieval process. In contrast, this study exploits 120 the soil moisture change (instead of the hydrometeors in the air) due to the recent rainfall impact 121 by using the non-raining observations at the low frequencies (19, 24, and 37 GHz) for daily rainfall 122 accumulation retrieval. It is worth mentioning that the non-raining observations account for $\sim 90\%$ 123 of the overall observations. 124

It is noted that Birman et al. (2015) used the surface emissivities at 89 GHz from multiple satellites to estimate the daily rainfall accumulation over France. As stated in the study, they use the "*effective emissivity that includes the atmospheric contribution in cases with cloudy/rainy conditions*", while this study only uses the non-raining emissivities for daily rainfall estimates at low frequency channels, including 19, 24 and 37 GHz. The data used in this study are described in section 2. The methodology, including the TB conversion, the land surface emissivity computation, the rainfall detection, and the daily rainfall amount estimation, are provided in section 3. Section 4 presents the retrieval results from the microwave emissivity temporal variation. Finally, the conclusions are summarized in section 5.

134 **2. Data**

This study primarily uses three types of datasets: the satellite TB observations, the IMERG final-135 run rainfall product, and the GPM Ku-band Precipitation Radar (KuPR) rainfall observations. 136 TB observations are from 10 passive microwave radiometers, including three SSMIS onboard 137 the Defense Meteorological Satellite Program (DMSP) F16, F17 and F18 satellites, AMSR2 on-138 board the Global Change Observation Mission-Water (GCOM-W) satellite, GMI onboard the 139 GPM core observatory satellite, four AMSU-A onboard NOAA-18, NOAA19, MetOp-A and 140 MetOp-B satellites, and ATMS onboard the Suomi National Polar-orbiting Partnership (SNPP) 141 satellite. The channels used in study and their mean footprint resolution are listed in Table 1. 142 These channels often have different footprint resolutions. Section 3 introduces a method to bring 143 these channels to a common resolution. Additionally, the frequencies among these satellites are 144 not identical (e.g., 18.7 v.s. 19.4 GHz). The slight frequency difference results in different TBs 145 for the same surface background and hydrometeor profile (Yang et al. 2014). We demonstrate a 146 method to convert TBs from other nine sensors to GMI frequencies in section 3. For convenience, 147 we do not distinguish the slight frequency differences among these sensors from now on unless 148 otherwise specified. These channels are referred to as V19, H19, V24, V37, H37, V89 and H89. 149 To derive the land surface emissivity at these frequencies, we use hourly surface temperature and 150 3-hourly temperature and humidity profiles at the 0.5°×0.625° resolution from Modern-Era Ret-151

rospective analysis for Research and Applications, Version 2 (MERRA-2) (Gelaro et al. 2017). We

also use the NOAA National Environmental Satellite, Data, and Information Service (NESDIS)'s
 daily Global 4km Multisensor Automated Snow/Ice map (GMASI) to filter out pixels associated
 with the snow/ice on the ground (Romanov 2017).

The half-hour IMERG final run (version 06A) product at 0.1 degree is used to investigate the 156 rainfall impact duration period. The IMERG final run daily dataset at 0.1 degree spatial resolution 157 is taken as the surface "reference" rainfall dataset for retrieval. In addition, we use the KuPR 158 rainfall observations along with the GMI TB observations to train a rain/no-rain screening method, 159 which is applied to all TB observations to filter out the raining pixels (more details in section 3c). 160 Data in this study are all from March 2014 (launch of the GPM satellite) to December 2018 over 161 the 60° S- 60° N land areas. The pixel level observations (TB and KuPR) are brought to a nominal 162 resolution of 59 km (see section 3 for more details). The IMERG data are downgraded to the 0.5° 163 spatial resolution by the simple arithmetic average. 164

3. Methodology

This section first describes a method to bring all channels from all sensors to a common nominal 166 resolution of 59 km. Then we discuss how to use Simultaneous Conical Overpass (SCO) technique 167 and Principal Component Analysis (PCA) for TB conversion (Yang et al. 2011; You et al. 2017a, 168 2018), where GMI is taken as the reference. Further, we briefly discuss the linear discriminant 169 analysis (LDA) approach for rain/no-rain screening (Turk et al. 2014; You et al. 2016, 2017b). Af-170 ter filtering out the raining pixels, we outline the procedure to compute the land surface emissivity 171 from TB observations for the non-raining pixels only (Munchak et al. 2020). Finally, we define 172 the emissivity temporal variation at the 0.5° resolution. 173

a. Collocation Scheme

The mean footprint resolution from these 10 satellites varies greatly, as shown in Table 1. This study takes the 19 GHz of SSMIS as the "nominal" resolution. The much finer spatial resolution footprints (e.g., 5 km) are averaged (downgraded) to this coarser resolution. The footprints with coarser resolution (e.g., 75 km) and resolution close to 59 km (i.e., 48 km) remain unchanged.

Specifically, for SSMIS we average 18 ($59 \times 59/14/14 \approx 18$) pixels of 85.5 GHz and 3 pixels of 179 37.0 GHz to match this nominal resolution. The resolution at 21.3 GHz is kept the same. For 180 GMI, we average 16 ($59 \times 59/15/15 \approx 16$) pixels of 18.7 GHz, 21 pixels of 23.8 GHz, 25 pixels of 181 36.6 GHz, and 71 pixels of 89 GHz to approximately match the 59 km resolution. For AMSR2, 182 we average 8 pixels of 18.7 GHz, 5 pixels of 23.8 GHz, 25 pixels of 36.5 GHz, and 140 pixels of 183 89.0 GHz to match the nominal resolution. For both ATMS and AMSU-A, the resolution at 23.8 184 and 31.4 remains unchanged. We average 4 pixels of 88.2 GHz from ATMS and 14 pixels of 89.0 185 GHz from AMSU-A to match the nominal resolution of 59 km. 186

To develop a rain/no-rain screening method, we use the KuPR observations at \sim 5 km resolution. After downgrading the GMI resolution to 59 km, we average 140 pixels of KuPR for each GMI pixel in GMI-KuPR overlapped region to match the nominal resolution.

For the MERRA2 ancillary variables, we use the nearest grid to match the nominal resolution at the closest time. In addition, a satellite pixel is judged as the snow/ice pixel if any of 4km GMASI grid is snow/ice-covered in the nominal resolution footprint, which is omitted in this study.

¹⁹³ b. Convert TB from other nine sensors to TB at the GMI frequencies

As shown in Table 1, the frequencies among these 10 radiometers are not identical. This study is to estimate rainfall accumulation by emissivity temporal variation derived from these TB observations. To this end, it is necessary to convert all TBs at similar frequencies to the same frequency. ¹⁹⁷ The conversion process has been detailed in You et al. (2017a, 2018). Here, only a brief summary ¹⁹⁸ is provided.

The following discussion takes the GMI and SSMIS (F18) as an example to discuss the conver-199 sion process, which can be summarized into four steps: (1) Find the simultaneous conical overpass 200 (SCO) pairs between GMI and SSMIS (Yang et al. 2011; You et al. 2017a, 2018). The SCO pairs 201 are pixels from these two satellites, which are at most 5 km apart and 5 minute away; (2) Decom-202 pose the GMI TBs from these SCO pairs into Principal Components (PCs); (3) Use the SSMIS 203 TBs in these SCO pairs to estimate the first several PCs by a linear regression model. This study 204 selects the first four PCs, which accounts for over 90% of the total variance; (4) Apply the coeffi-205 cients derived from the SCO pairs to the whole SSMIS data. By doing so, we obtain the estimated 206 PCs from SSMIS. These PCs are converted back to TBs at the GMI frequencies. 207

The same procedure is applied to AMSR2, SSMIS (F16 and F17), ATMS and AMSU-A. The 208 V37 GHz channel from F17 is not used since the data from April 2016 are not processed by the 209 calibration team due to the large noise. The missing V37 channel on F17 SSMIS shows little 210 influence on the TB conversion. On the other hand, both ATMS and AMSU-A only have the 211 vertically polarized channels, and both radiometers do not have channels around 19 GHz. Later 212 analyses will show that the root-mean-square-error (RMSE) from the TB conversion based on 213 ATMS and AMSU-A is noticeably larger than those from AMSR2 and SSMIS. However, section 214 4 clearly demonstrates the improved rainfall retrieval performance by including these five sounders 215 due to the increased sample size. 216

In contrast to our previous studies (You et al. 2017a, 2018), this study applies the TB conversion procedure at each 2.5° grid box. By doing so, we show later that RMSE from this conversion is less than 3 K over almost all the areas from 60°S to 60°N. After this TB conversion process, GMI and other nine sensors all have channels of V19, H19, V24, V37, H37, V89 and H89. The objective of this study is to use the emissivity under non-raining conditions to retrieve daily rainfall accumulation. To this end, we use the linear discriminant analysis (LDA) approach (Turk et al. 2014; You et al. 2015) to filter out the raining pixels. This method is first developed based on GMI and KuPR observations, then applied to converted TBs from other nine sensors.

Suppose there exist two training databases from KuPR (i.e., raining vs. non-raining databases), which contain multi-variables \mathbf{x} (i.e., V19,..., H89) in each database. According to Wilks (2011) the linear discriminant function to distinguish these two groups is:

$$\boldsymbol{\delta}_1 = \boldsymbol{a}^T \times \mathbf{x} \tag{1}$$

Where T stands for the transpose. a is the discriminant vector, calculated in the following way:

$$\boldsymbol{a} = \mathbf{S}_{pool}^{-1}(\bar{\mathbf{x}}_1 - \bar{\mathbf{x}}_2)$$
$$\mathbf{S}_{pool} = \frac{n_1 - 1}{n_1 + n_2 - 2} \mathbf{S}_1 + \frac{n_2 - 1}{n_1 + n_2 - 2} \mathbf{S}_2$$
(2)

 $\bar{\mathbf{x}}_i$ and \mathbf{S}_i (i = 1, 2) represent the mean vector and covariance of each group, respectively. \mathbf{S}_{pool} is the weighted average of the two sample covariance matrices from these two datasets. n_1 and n_2 are the samples size in these two groups, respectively.

233 *d. Emissivity computation*

We compute the emissivity values for each pixel at different channels based on Munchak et al. (2020), which is briefly summarized here. The emissivity vector is calculated from the converted brightness temperatures (i.e., other satellites' observations are converted to the GMI channel set). This allows us to use the same atmospheric absorption and incidence angle assumptions that are used for GMI in Munchak et al. (2020). The emissivity and atmospheric temperature and water vapor profile are retrieved using an optimal estimation inversion procedure. For the set of channels
used in this study, however, there is little independent information about the atmospheric profile,
and the retrieved emissivities are essentially those that reproduce the converted brightness temperatures, given the space-time interpolated MERRA2 skin temperature and atmospheric profile.

243 e. Emissivity temporal variation definition

To derive emissivity temporal variation, it is necessary to determine when the observations from different satellites are considered as observations for the same location. This study first divides the globe into a 0.5° grid box. We define any observation in the same 0.5° latitude-longitude grid box as observations at the same location. We choose the 0.5° grid box because the nominal resolution (59 km) is approximately 0.5° in the tropical region. Choosing a different grid size (e.g., 0.25° or 1°) does not affect the major conclusions of this work.

The emissivity (e) temporal variation (Δe) is defined as:

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$$\Delta e = e_{t_0} - e_{t_{-1}} \tag{3}$$

$$\Delta t = t_0 - t_{-1} \tag{4}$$

²⁵³ Where e_{t_0} is the current daily mean emissivity when rainfall occurs, and $e_{t_{-1}}$ is the preceding ²⁵⁴ daily mean emissivity at the same location without rainfall. The daily emissivity is computed as ²⁵⁵ the arithmetic mean from the selected satellites (e.g., all-10-satellites or imagers-only, see section ²⁶⁶ 4e for details). A day is judged as raining day when there is at least one raining observation on that ²⁶⁷ day. No raining pixels are included in the daily mean emissivity average process. Δt is the time ²⁶⁸ difference between these two days. From now on, the Δe at V19, H19, ..., V89 will be referred to ²⁵⁹ as Δe_{v19} , Δe_{h19} , ..., Δe_{v89} for convenience.

260 f. Bayesian retrieval method

To retrieve the daily rainfall accumulation from Δe , we adopt the Bayesian retrieval technique as implemented by You et al. (2016). It is worth mentioning that the Bayesian retrieval concept is widely used in the precipitation/cloud retrieval community (e.g., Evans et al. 1995; Kummerow et al. 1996; Chiu and Petty 2006; Noh et al. 2006; Kim et al. 2008; Sano et al. 2013; Petty and Li 2013a; You et al. 2015, 2016).

²⁶⁶ Mathematically, the retrieval method can be stated as follows:

$$f(x|\mathbf{y}) = \frac{f(\mathbf{y}|x) \times f(x)}{f(\mathbf{y})}$$
$$= \frac{f(\mathbf{y}|x) \times f(x)}{\int f(\mathbf{y}|x) \times f(x) dx}$$
(5)

where *x* and *y* represent the daily rainfall amount and the emissivity temporal variation vector ([$\Delta e_{v19}, \Delta e_{h19}, \Delta e_{v24}, \Delta e_{v37}, \Delta e_{h37}$]), respectively. Later analyses will show that emissivity at 89 GHz has very weak response to the previous rainfall, compared with the low frequencies. Therefore, in the rainfall retrieval process, we only include the emissivity at 19, 24 and 37 GHz and emissivity at 89 GHz is not used. f(x|y) is the posterior probability density function (PDF) of *x* given the *y*, f(x) is the prior PDF of *x* and f(y|x) is the likelihood function of *y* given the precipitation rate *x*.

The expected value of x is taken as the final estimation for the daily rainfall amount, which is computed in the following way:

$$E(x|\mathbf{y}) = \frac{\int x \times f(\mathbf{y}|x) \times f(x)dx}{\int f(\mathbf{y}|x) \times f(x)dx}$$
$$= \frac{E[x \times f(\mathbf{y}|x)]}{E[f(\mathbf{y}|x)]}$$
(6)

where E stands for the expectation.

277 **4. Results**

This section first shows the TB conversion and the rain/no-rain detection statistics. Then, we explain how to determine the rain-sensitive-region based on the emissivity depression corresponding to the different daily rainfall amount. We also explain why we would like to retrieve the daily rainfall amount, instead of multiple-day rainfall accumulation. Finally, we present the retrieval results from four different satellite constellation experiments and demonstrate why satellites with varying equator crossing times are necessary for the best retrieval performance.

²⁸⁴ a. Brightness Temperature Conversion Statistics

Figure 1 shows the sample size of the SCO pairs in each 2.5° grid box between GMI and other nine sensors over land. It is found that the sample size in the vast majority of boxes (>99%) for all satellites is greater than 200, which is sufficient to ensure the conversion coefficients are stable.. In case there are not enough SCO pairs (<200) in some grid boxes, especially from MetOp-A (Fig. 1f) and MetOp-B (Fig. 1g), we aggregate the SCO pairs in the nearest several grid boxes until the sample size is greater than 200.

Figure 2 shows a conversion case study at H19, H37, and H89 at the grid box of (32.5°N, 103°W) 29 between GMI and AMSR2 (Fig. 2a to Fig. 2c), between GMI and SSMIS-F18 (Fig. 2d to Fig. 2f), 292 and between GMI and ATMS (Fig. 2g to Fig. 2h). The plots from SSMIS-F16 and SSMIS-F17 are 293 similar to those from SSMIS-F18. The plots from four AMSU-A sensors are similar to those from 294 ATMS. It is noticed that the estimated GMI TBs from these three satellites are very close to GMI 295 observations. In fact, the correlation from all these channels are over 0.95, and the bias is close to 296 0, which indicates that the conversion is working correctly. The RMSE (shown on the figure) is 297 less than 3 K, except the estimated H19 from ATMS. As mentioned earlier, ATMS does not have 298 frequency around 19 GHz. Also, only the vertically polarized channels are available from ATMS 299

(Table 1). These two ATMS features are responsible for the larger RMSE from ATMS. Similarly, RMSE from H19 estimated from AMSU-A is also noticeably larger than those from AMSR2 and SSMIS. Further, RMSE at H19 and H37 from AMSR2 is smaller than those from SSMIS, which is likely due to the finer footprint resolution from AMSR2, and the almost identical frequencies between GMI and AMSR2. At 89 GHz, RMSE from AMSR2 (Fig. 2c) and SSMIS-F18 (Fig. 2f) is comparable, likely due to the large impact of the hydrometeors in the atmosphere.

The RMSE global distribution at H19, H37, and H89 is shown in Fig. 3 for AMSR2, SSMIS-306 F18, and ATMS. Our analysis shows that over 95% of the grid has a RMSE less than 3 K, which 307 corresponds to ~ 0.01 emissivity error. Consistent with the case study, RMSE from ATMS is the 308 largest in almost all regions. RMSE at H19 and H37 GHz from AMSR2 is noticeably smaller than 309 those from SSMIS (cf. Fig. 3a and Fig. 3d, cf. Fig. 3b and Fig. 3e). For H89 channel, RMSE 310 from SSMIS and AMSR2 are of comparable magnitude. Analysis has also been conducted for all 31 the vertically polarized channels (V19, V24, V37, and V89), yielding very similar results to those 312 from the horizontally polarized channels. 313

314 b. Rainfall Detection Statistics

Similar to the TB conversion process, we refine our previously developed LDA rainfall detection method by applying it to each 2.5° grid box. To ensure the stability of the detection statistics, the number of raining pixels in each 2.5° grid box is required to be at least 500. When there are less than 500 raining pixels, we aggregate the observations in the nearest several 2.5° grid boxes until the sample size is greater than or equal to 500. At each grid box, a discriminant threshold value is selected to maximize the Heidke Skill Score (HSS).

Figure 4 shows that the Probability of Detection (POD) and HSS are over 0.7 over the majority of the region, and the False Alarm Rate (FAR) is less than 0.05 over most of the region. These detection statistics are similar to those from the official National Aeronautics and Space Administration (NASA) and the Japan Aerospace Exploration Agency (JAXA)'s precipitation detection algorithms (You et al. 2020).

We would like to emphasize that in the daily rainfall estimation process, we first filter out the raining pixels judged by the LDA detection method. Therefore, the signal we use is essentially the soil emission variation due to the rainfall impact, not the hydrometeors' effect in the air.

329 c. Rainfall Sensitive Regions

Previous studies (Brocca et al. 2014; McColl et al. 2017) showed that rainfall has little impact on the land surface soil moisture derived from microwave radiometer observations at 1.5 GHz over densely vegetated regions (e.g., Amazon, Central Africa, and Eastern United States). This study primarily exploits the soil moisture change due to the recent rainfall impact at the low frequency channels. Therefore, we would like first to select rainfall sensitive regions, by assessing the surface emissivity response to rainfall over different regions. This analysis is based on the emissivity derived from GMI only to reduce the computational time, instead of from all 10 satellites.

Figure 5 shows the emissivity depression at H19 and H37 corresponding to different previous 1-day rainfall accumulation. Specifically, we obtain the previous 1-day rainfall amount corresponding to each pixel from the half-hour IMERG Final-run product. Then, we compute the emissivity differences between wet (rainfall occurs in the previous one day) and dry (no rainfall in the previous one day) conditions at the 0.5° resolution. For the wet condition, the previous 1-day rainfall accumulation (indicated by R on the Fig. 5) is further grouped into four categories, including 0 < R < 5, 5 < R < 10, 10 < R < 20, and R > 20 mm.

As illustrated in Fig. 5, emissivity decreases over most of the land areas after rainfall events in the previous day, and the emissivity depression increases as the rainfall amount becomes larger.

The emissivity drop is particularly evident with rainfall accumulation greater than 20 mm over Sa-346 hel, Southern Africa, Middle East, Indian sub-continent, northwest China, Australia continent, and 347 western United States (Fig. 5d and Fig. 5h). As expected, the emissivity depression magnitude 348 is smaller at H37 than at H19 since 19 GHz is more sensitive to the surface properties (e.g., soil 349 moisture). The emissivity depression at H89 (not shown) is even smaller than H37. Similar analy-350 sis has also been performed for the vertical polarized channels, from which the rainfall response is 35 weaker than their horizontally polarized counterparts. These features (e.g., lower frequency with 352 larger emissivity drop due to the rainfall impact) are well-known from previous studies (Jackson 353 1993; Ferraro et al. 2013; You et al. 2014; Munchak et al. 2020). Based on these analyses, we only 354 use the emissivity values at V19, H19, V24, V37 and H37 for the daily rainfall retrieval and no 355 emissivity values from 89 GHz are included in the retrieval process. 356

This study attempts to exploit the emissivity depression signature due to the recent rainfall impact. For this purpose, we define regions with emissivity drop of at least 0.02 with previous 1-day rainfall accumulation greater than 20 mm as "rainfall-sensitive-regions" and retrieval is only performed over these regions.

d. Correlation between emissivity and rainfall accumulation at different time scales

It is desirable to understand how long the rainfall impact can last. To this end, we compute the correlation between emissivity under rain-free conditions and previous n-hour rainfall accumulation at the 0.5° resolution over the rainfall-sensitive-regions. Similar to the rainfall-sensitive analysis, this analysis also only uses the emissivity derived from GMI to reduce the computational time.

³⁶⁷ We first attach the previous rainfall accumulations of 1-hr, 3-hr, 6-hr, 12-hr and 24-hr to each ³⁶⁸ emissivity pixel. Then, we compute the correlation between emissivity and these rainfall accumulations at each 0.5° grid box. The purpose is to check when the correlation peaks. Fig. 6 shows that
the correlation magnitude between emissivity and rainfall accumulation increases quickly from 1hr (Fig. 6a) to 12-hr (Fig. 6d) for H19, which is especially evident over Australia. While for the
correlation from 12-hr (Fig. 6d) to 24-hr (Fig. 6e), such an increase is marginal. The H37 channel
exhibits similar correlation temporal variations (Fig. 6f to Fig. 6j).

To more clearly show this correlation variation, we plot the histogram of the correlation be-374 tween emissivity and previous N-hr rainfall accumulation (N varies from 1 to 24-hr). The curves 375 corresponding to 12-hr (black) and 24-hr (purple) are heavily overlapped (Fig. 7), demonstrating 376 that the correlations between emissivity and 12-hr rainfall accumulation is very similar to the cor-377 relation between emissivity and 24-hr rainfall accumulation. This implies that previous rainfall 378 being 13 to 24-hr away from that pixel has little impact to the emissivity value of that pixel. In 379 other words, rainfall impact at H19 and H37 often persists about 12-hr. Therefore, this behavior 380 supports our goal to estimate the daily rainfall accumulation. Of course, the rainfall impact can 38 propagate into the next day, depending on at which hour the rainfall occurs. While 2-day or 3-day 382 accumulations might be more accurate, most applications prefer daily accumulations. 383

384 e. Rainfall retrieval result

This section begins with a case study to explain in detail how the retrieval algorithm is implemented. Then we test four retrieval experiments to show why using 10 satellites produces the most accurate retrieval result.

Figure 8a shows the time series of emissivity at H19 over a selected grid box (32°N-32.5°N, 99.5°W-100°W) in the Central Texas region. The LDA detection approach indicates that rainfall occurs on 23, 24, and 25 May 2014 (shown as red crosses). On 24 and 25 May 2014, these rainfall occurrences correspond very well with the "cold" TB at V89, shown in Fig. 8b. On 23 May 2014, the TB depression at V89 is not as obvious as that on 24 and 25 May 2014. However, IMERG rainfall product indeed shows the daily rainfall amount at 1.2 mm on 23 May 2014.

The daily rainfall retrieval algorithm takes the following steps: (1) filter out the raining pixels 394 (indicated by the red crosses) from 23 May to 25 May 2014 because the computed emissivities 395 for these pixels are affected by the hydrometeors in the atmosphere; (2) compute the daily mean 396 emissivity using the non-raining values; (3) compute the daily mean emissivity on the preceding 397 non-raining day (i.e., 22 May 2014, background emissivity); (4) obtain the emissivity difference 398 between raining day (23, 24, and 25 May 2014) and background emissivity (22 May 2014). For 399 demonstration purposes, Fig. 8a only shows the emissivity temporal variation at H19. In the 400 retrieval process, we use the emissivity variation at 19, 24 and 37 GHz (i.e., Δe_{h19} , Δe_{v19} , Δe_{v24} , 401 $\Delta e_{\nu 37}, \Delta e_{h 37}$). This Δe computation procedure is applied at each 0.5° grid box over the rainfall-402 sensitive regions. 403

Next, we design four experiments to demonstrate the advantages of using multiple satellites. In 404 each experiment, we randomly select 80% data at each grid box as the training dataset, while the 405 retrieval is performed on the other 20% data. In the first experiment, Δe at each channel is calcu-406 lated by GMI observations only. The second experiment computes Δe using all five imagers (four 407 sun-synchronous satellites shown in blue color in Fig. 9), including GMI, AMSR2, and three 408 SSMISs. Clearly, the Equator Crossing Times (ECTs) from F17 and F18 are, on average, only 409 about 10 minutes apart, indicating that they observe the same location at nearly the same time of 410 day. In the third experiments, we select six sensors, including GMI and five other radiometers on-411 board the sun-synchronous satellites (i.e., AMSR2 onboard GCOM, AMSU-A onboard NOAA19, 412 SSMIS onboard F16 and F17, and AMSU-A onboard Metop-A). The selection of these five sun-413 synchronous sensors is based on the fact that ECTs from them are very different, as shown in Fig. 414 9. By doing so, the emissivity temporal variation can be better captured. The fourth experiments 415

use observations from all 10 sensors to compute Δe . For convenience, these four experiments are referred as "GMI only", "5-imagers", "6-satellites", and "all-10-satellites".

When only the GMI is used to compute Δe , the retrieval performance is rather poor, as indicated 418 by the correlation being 0.25 and RMSE being 11.36 mm (Fig. 10a). It is immediately clear 419 that the 5-imagers scheme produces much improved retrieval results. Specifically, the correlation 420 increases to 0.49 and RMSE decreases to 7.28 mm (Fig. 10b). Further analysis reveals several 421 reasons responsible for this large retrieval improvement, which are all related to the observation 422 sample size. First, the time difference (i.e., Δt in Eq. 2, the time difference between the raining 423 day and the non-raining day) is shorter when using five imagers than only using GMI, as shown 424 in Fig. 11. The time difference is one-day for over 85% cases when using five imagers, which 425 means that one can find a non-raining background in the preceding day when using five imagers 426 for over 85% of the time. In contrast, only about 34% of the time one can find a non-raining 427 background when only GMI observations are used. For the majority of the time, the non-raining 428 background is two, three, or even more days away when only GMI is used. With the longer time 429 difference, it is more likely that the emissivity varies due to factors other than the rainfall impact, 430 or the rainfall effect might be missed. Second, with more observations from five satellites, the 431 diurnal cycle of the emissivity can be much better captured than that using GMI observations only. 432 In fact, on average, the daily sample size over each 0.5° grid box is 10 when using 5 imagers, 433 while it is only 1 or 2 from GMI. Lin and Minnis (2000) found that the emissivity of 19 GHz from 434 Special Sensor Microwave Imager (SSM/I) at the early morning (06:40 local time) is about 0.06 435 less than that at other times over a Southern Great Plains site, and they concluded that dew and 436 surface rewetting effects may be responsible for the emissivity diurnal cycle. The large emissivity 437 discrepancy between daytime and nighttime (up to 0.1 over some arid regions) has also been 438 noticed by Norouzi et al. (2012) using AMSR-E observations, although they pointed out that the 439

different diurnal cycles between the skin temperature and the soil temperature are responsible for
the large emissivity discrepancy. Regardless of the underlying mechanisms causing the emissivity
diurnal cycle, more observations from multiple satellites can better capture the daily emissivity
variation compared with those from a single satellite. Third, the increased temporal sampling
from multiple satellites provides a better chance of an observation right after the rainfall has ended,
when its effect on emissivity is maximum.

By carefully selecting six sensors with much different ECTs, the retrieval performance is further 446 improved, indicated by the correlation being 0.58 and RMSE being 6.99 mm (Fig. 10c). The time 447 difference between using five imagers and using six sensors is similar (Fig. 11). That is, over 85% 448 of the time difference in both experiments is one day. However, with the much variable ECTs from 449 the 6-satellites scheme, the emissivity variation can be better captured than that in the 5-imagers 450 scheme. As mentioned previously, ECTs from F17 and F18 are very similar from 2014 to 2018 45[.] (Fig. 9). By using observations from all-10-satellites scheme, the retrieval results only improve 452 marginally with the correlation being 0.60 and RMSE being 6.52 mm, compared with that from 453 the 6-satellites scheme. The marginal improvement is expected since ECTs from several satellite is 454 similar (Fig. 9, Metop-A and MetopB, F17, F18 and NOAA18, AMSR2 and ATMS). This means 455 that observations from these satellites with similar ECTs add little new information. 456

⁴⁵⁷ A common feature in the retrieval result from Fig. 10b to Fig. 10d is that for rain rates less than ⁴⁵⁸ 1 mm, the retrieval algorithm has little skill. This phenomenon may reflect the fact that the soil ⁴⁵⁹ moisture has little response for daily rainfall accumulations less than 1 mm.

5. Conclusions and Discussions

This study presents a rainfall retrieval algorithm to estimate the daily rainfall accumulation from non-raining satellite observations from 10 satellites, including GMI, AMSR2, SSMIS onboard F16, F17, and F18 satellites, ATMS onboard SNPP satellite, and AMSU-A onboard NOAA-18, NOAA-19, MetOp-A and MetOp-B satellites. In contrast to the traditionally used ice-scattering signal over land, we use the land surface emissivity variation signature due to the rainfall impact for rainfall retrieval by filtering out the raining pixels. To compute the emissivity temporal variation, we first convert frequencies from other sensors to GMI frequencies from 19 (or 24) to 89 GHz. Results show that RMSE is less than 3 K over the vast majority of the regions for all nine sensors and for all channels, leading to about 0.01 emissivity uncertainty.

The objective of this study is to use the non-raining pixels to compute the emissivity. To this end, we need to filter out the raining pixels first. Our statistical method shows strong capability to detect raining pixel, indicating by POD and HSS greater than 0.70 over the majority of the region. The rainfall retrieval algorithm is only applied to the rainfall-sensitive-region, defined as the areas where the land surface emissivity drops at least 0.02 at H19 corresponding to the previous 1-day rainfall accumulation greater than 20 mm.

While the best rainfall retrieval performance is achieved by using observations from all-10-476 satellites scheme, with the correlation and RMSE being 0.60 and 6.52 mm, analysis shows that by 477 selecting GMI and five sensors onboard the sun-synchronous satellites with much different ECTs 478 (i.e., 6-satellites scheme), the retrieval performance is comparable to that from 10 satellites, as 479 indicated by the correlation of 0.58 and RMSE of 6.99 mm. In contrast, the retrieval results from 480 the 5-imagers scheme are noticeably worse than those 6-satellites and all-10-satellites schemes 481 because the emissivity variation can be much better captured by using all 10 satellites or six satel-482 lites with much different ECTs, compared with only using five imagers. Furthermore, there is 483 low retrieval skill when only the GMI observations are used due to the much smaller sample size, 484 which leads to a longer time difference between the raining day and the non-raining day. Also, it 485 is not possible to capture the emissivity diurnal cycle with GMI observations only. 486

Future work seeks to further include the currently operational radiometers, including ATMS onboard NOAA-20, AMSU-A onboard Metop-C, WindSat, and FengYun-3 Microwave Radiometer Imager (MWRI). With more observations, the retrieval performance from our method is expected to be further improved. In particular, we expect a large retrieval performance improvement when the passive microwave radiometer observations around 8:00 am and 11:00 am are available (see Fig. 9) in the future.

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645	Table 1.	Channels used for rainfall retrieval from each sensor (V-Vertical polarization,
646		H-horizontal polarization). The sensors employed the cross-track scanning
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Satellite Name	Sensor Name	Freq (GHz)	Freq (GHz)	Freq (GHz)	Freq (GHz)
GPM	GMI	18.7 (V/H, 15 km)	23.8 (V, 13 km)	36.6 (V/H, 12 km)	89.0 (V/H, 7 km)
GCOM	AMSR2	18.7 (V/H, 22 km)	23.8 (V/H, 26 km)	36.5 (V/H, 12 km)	89.0 (V/H, 5 km)
F16	SSMIS	19.4 (V/H, 59 km)	21.3 (V, 59 km)	37.0 (V/H, 36 km)	85.5 (V/H, 14 km)
F17	SSMIS	19.4 (V/H, 59 km)	21.3 (V, 59 km)	37.0 (V/H, 36 km)	85.5 (V/H, 14 km)
F18	SSMIS	19.4 (V/H, 59 km)	21.3 (V, 59 km)	37.0 (V/H, 36 km)	85.5 (V/H, 14 km)
NPP	ATMS*		23.8 (V, 75 km)	31.4 (V, 75 km)	88.2 (V, 32 km)
NOAA-18	AMSU-A*		23.8 (V, 48 km)	31.4 (V, 48 km)	89.0 (V, 16 km)
NOAA-19	AMSU-A*		23.8 (V, 48 km)	31.4 (V, 48 km)	89.0 (V, 16 km)
MetOp-A	AMSU-A*		23.8 (V, 48 km)	31.4 (V, 48 km)	89.0 (V, 16 km)
MetOp-B	AMSU-A*		23.8 (V, 48 km)	31.4 (V, 48 km)	89.0 (V, 16 km)

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655 656 657 658 659	Fig. 1.	Coincident observation number in each 2.5° grid box between GMI and other 9 sensors, in- cluding AMSR2, SSMIS-F16, SSMIS-F17, SSMIS-F18, ATMS-SNPP, AMUSA-MetOpA, AMSUA-MetOpB, AMSUA -NOAA18, and AMSUA -NOAA19. The number is scaled by 100 in each plot. All data are from March 2014 (launch of the GPM satellite) to December 2018	б
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698	Fig. 10.	(a) Density scatter plot between IMERG daily rainfall amount and retrieved daily rainfall	
699		amount based on the emissivity temporal variation (Δe) at 19, 24 and 37 GHz, derived from	
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701		including GMI, AMSR2, and three SSMIS. (c) Same as (a) except the Δe is derived from six	
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707		Eq. 2, for all four retrieval experiments shown in Fig. 10.	5



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